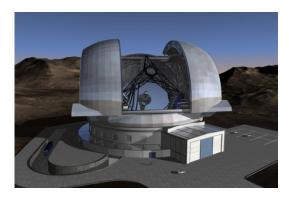
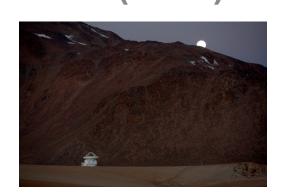




Challenges in Modern Astrophysics: Ground-based facilities



Bruno Leibundgut (ESO)







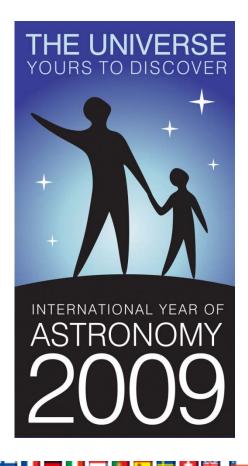
Astrophysics in a Golden Age

- Full coverage of electro-magnetic spectrum
 - MAGIC/HESS → Fermi/INTEGRAL → XMM/Chandra/Swift/Rossi
 XTE → Galex → HST/Gaia → ground-based optical/IR → Spitzer
 → Herschel → Planck → IRAM/JCMT/APEX/ALMA → radio
 telescopes
 - 20 orders of magnitude in wavelength/frequency/energy
 - Large archive collections (e.g. ROSAT, ISO, ESO, HST, MAST)
- Astro-particles joining in
 - cosmic rays, neutrinos, gravitational waves, dark matter searches



Astrophysics in a Golden Age

- International Year of Astronomy
 - Fantastic boost in the public
 - Increased awareness
 - Strong public support
 - Continued interest
 - Connected to the 'big' questions
 - Where do we come from?
 - What is our future?





Fantastic opportunities

Already existing ground-based facilities in Europe

Westerbork, Roque de los Muchachos (GTC, WHT, TNG, NOT, ING, MAGIC), Solar telescope on El Teide, Effelsberg, JCMT, La Silla, Paranal, IRAM (Plateau de Bure, Pico Veleta), HESS

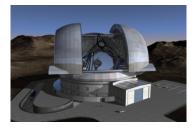
New facilities

VISTA, VST, LOFAR, ALMA

Under discussion E-ELT, SKA, EST, CTA









European ground-based telescopes

- Spain
 - GTC, WHT, ING, TNG, NOT, MAGIC, Calar Alto, IRAM
 - Solar telescopes
- France
 - CFHT, (Dome C on Antarctica), IRAM
- Germany
 - LBT, HET, SALT, Calar Alto, MAGIC, HESS, APEX, IRAM
- United Kingdom
 - Gemini (North and South), (AAT), UKIRT, WHT, JCMT, eMERLIN
- Italy
 - LBT, TNG, (Dome C)
- International
 - European VLBI Network, LOFAR, ESO

Access to telescopes

- Most observatories have open access policies
 - Preference for supporting institutes
- OPTICON Access program
 - Access to 4m telescopes is supported by OPTICON funding for successful proposals
 - ESO 3.6m, NTT, Calar Alto, WHT, TNG, and various others
- Archives!



You do not have to observe yourself any longer!

- Service observing at many telescopes
- Massive surveys publicly available
 - Sloan Digital Sky Survey (SDSS)
 - CFHT Legacy Surveys
 - UK Infrared Deep Sky Survey (UKIDSS)
 - GOODS/COSMOS
 - VST Public Surveys
 - VISTA Public Surveys
 - PanSTARRS (US)
 - Dark Energy Survey (US/European)
 - Large Synoptic Survey Telescope (LSST US; proposed)
- Large Archives



ING, ESO, HST/MAST (US and mostly space)

Presentation by Eduardo Gonzalez



Science themes

- What matters in the universe?
- Planets, planets, planets
- How did stars and planets form?
- The Milky Way our Home
- Our own black hole
- How galaxies form and evolve?
- Fashions and other transients
- When opportunity knocks



What matters in the Usiver

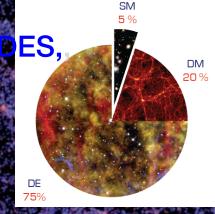
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– Requires for samples – Middeyear and (often) moltreelesso joe

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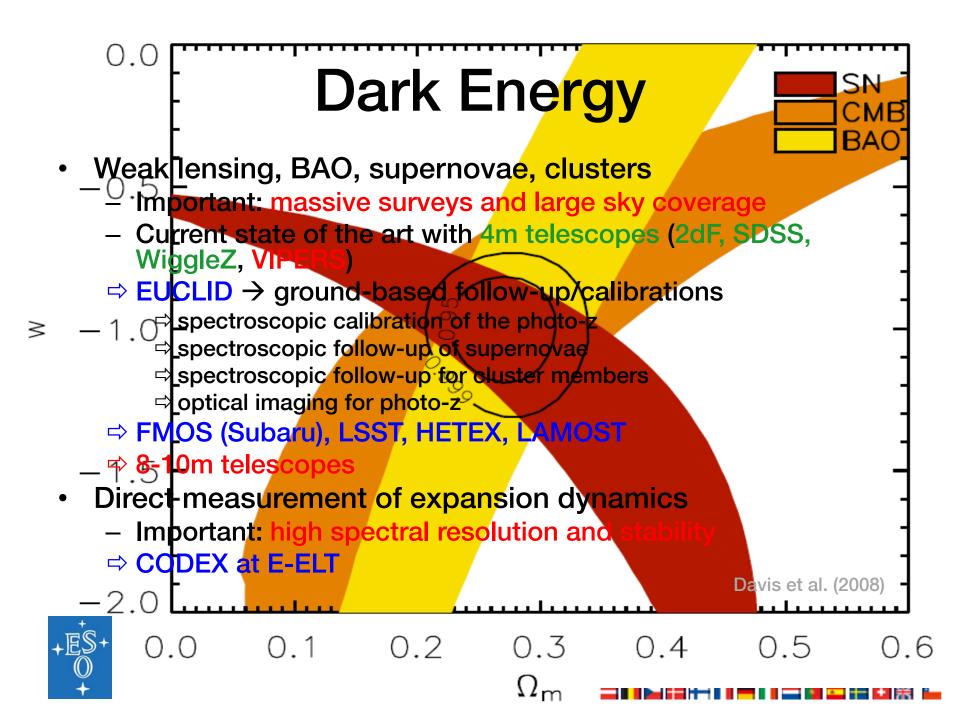
 Celoxy clusters (REFUE Nonae SPT DES, cROSPLA LOST)
 Hedishi^{er} distortions (VVDS, VIPER)

Springel

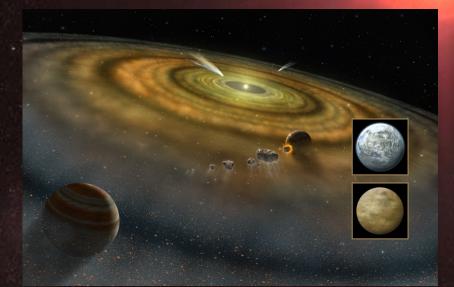


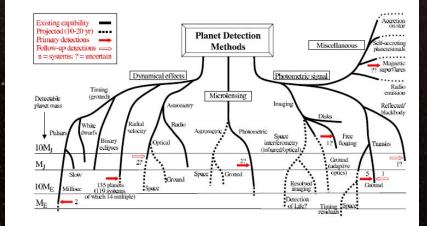


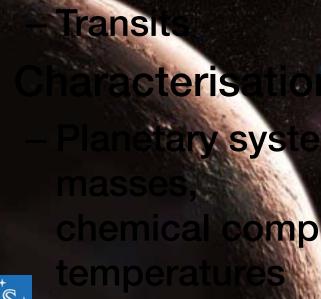
Millenium



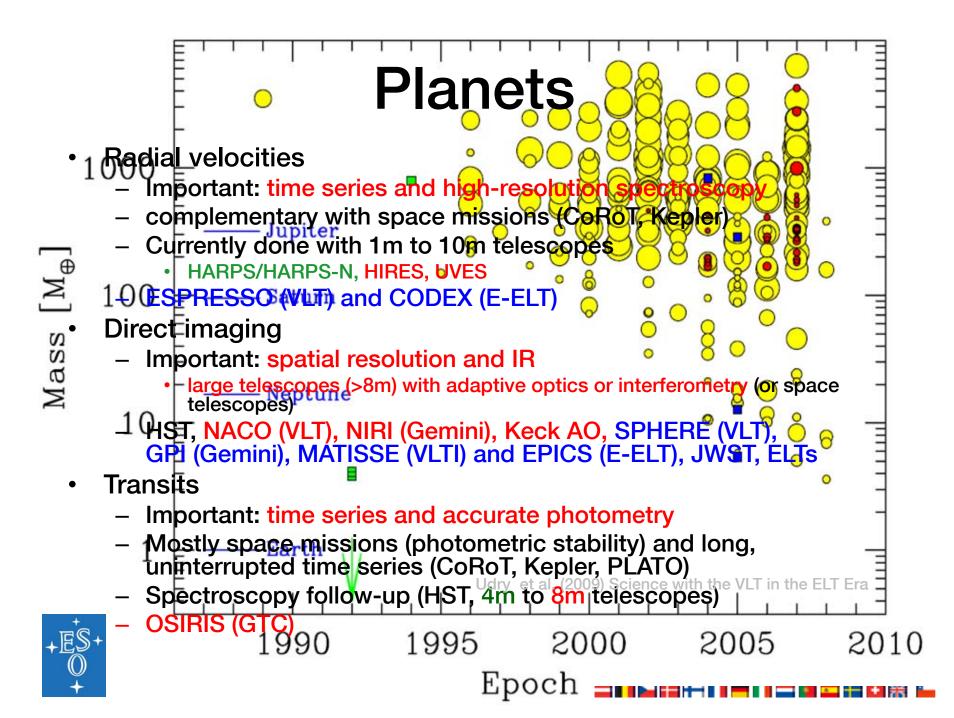
nets, planets











The ESO exo-planet machinery

• HARPS at 3.6m telescope

- best radial velocity machine at a 4m telescope (supported by UVES on VLT)
- extremely stable spectrograph
- fast pipeline \rightarrow nearly immediate results
- NACO
 - adaptive optics system on an 8m telescope
- VLTI
 - highest spatial resolution for follow-up observations of known systems
- NACO/SINFÓNI/FORS2
 - transit measurements
 - atmospheres of exo-planets



Searching for other earths

- HARPS is the most successful planet finder right now
- Almost all low-mass planets found with







Gamma-Ray Bursts

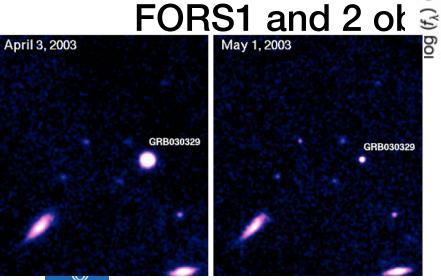
- Identification relied on optical data

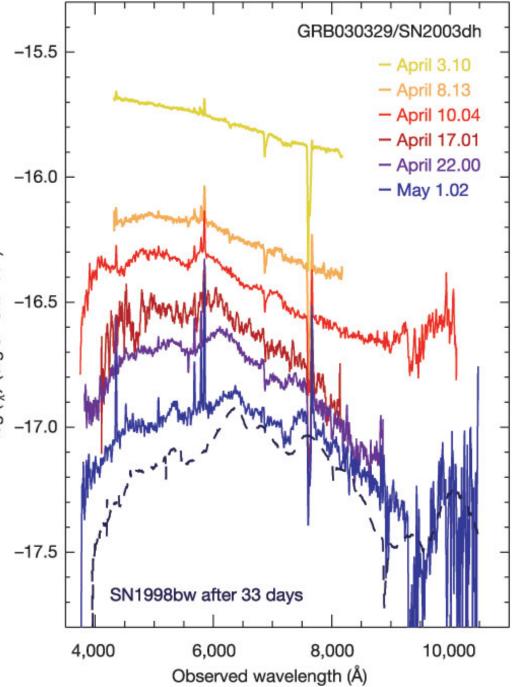
 redshifts, explosion energies, explosion physics
- Cosmological probes
 - the most distant observable stars
 - light houses to measure the intergalactic medium
 - tracers of chemical enrichment?
- Very short duration
 - required special instrumentation and software to observe adequately



SN/GRI

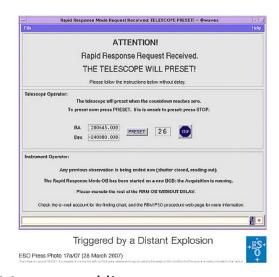
- Spectral signation appear in Gam
- GRB 030329/S GRB 980425/S UVES spectrum closest knowr FORS1 and 2 ot





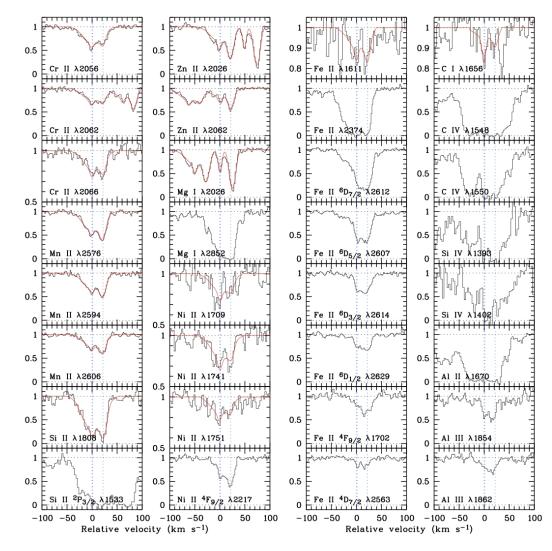
Rapid Response Mode (RRM)

UVES observations of GRB 060418 10 minutes after the initial Swift trigger



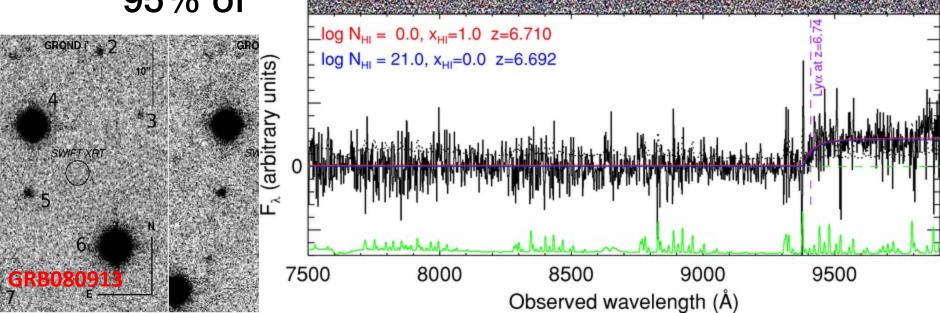


Many metal line systems at 3 redshifts. [Zn/Fe] >> QSO abs.



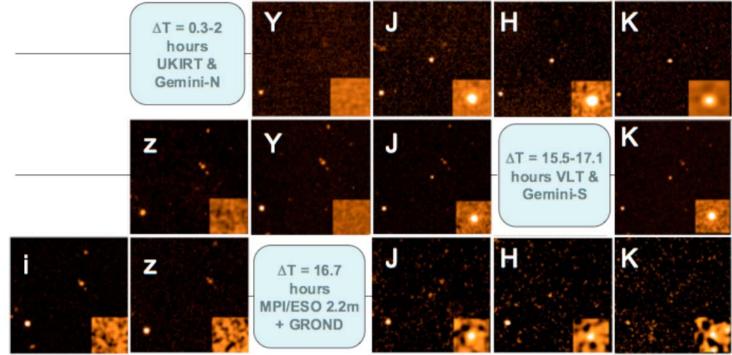
Gamma-Ray Bursts

- Most distant stellar objects ever observed
 - redshifts 6.7 and 8.2 (tentative)
 - lookbac⁻ 95% of ⁻



Most distant stellar object yet observed – GRB 090423

- Optical drop-out, bright in the near-infrared
- Rapid decline

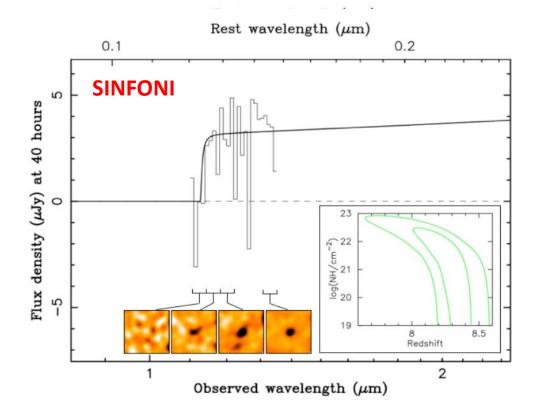


Tanvir et al., Nature submitted



GRB 090423

- Spectroscopy 17 hours after outburst
- Lyman break indicates a redshift of z≈8.2





Star and planet formation

Observing the warm cores of molecular

Implementation spatial resolution and large wavelength coverage
 R observations with large (>8m) telescopes,
 CanariCam (GTC), VLTI (MATISSE JWST, ELTs
 ALMA will be the champion this field



Wolf & Klahr 2005

Radial veloci u Cep lecades years 15 7234 y and Hipparcos parallaxes Spiral arms Gas flows, stellar distribution **Bulge composition, Galactic Centre** Distribution of massive stars 8 CMa

Sag

o Cas



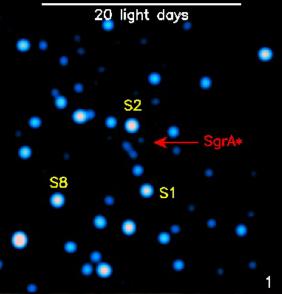
The M

Our own black hole

Mass determination through stellar orbits Structure around the black hole revealed through flashes **Coordinated studies** with other wavelengths







- 0.4 Galactic Centres31 519
- Determine the black hole event horizon
 Schwarzschild radius 9 microarcseconds
- Measure gravity in the strong regime
 Probing the spacetime geometry 538 521
 - Important: IR observations and spatial resolution ->slarge telescope (\$8m) with AO and interferometry 533

- NACO, Keck AO, GEMS (Gemini), GRAVITY (VLTI), ELTS 0.4 0.2 0. -0.2 -0.4 B.A. (")



How did galaxies form and evolve?

- Characterisation of the Lyman-break galaxies
 - Galaxy population at z>3
- Discovery of compact, old galaxies at z>1
 - "red and dead", "red distant galaxies"
 - Characterisation of galaxies at high z
 - Internal kinematics
- Earliest observable stellar agglomerations
 - Ly-a emitters





The distant universe

- Build up of the Hubble sequence
 - Star forming vs. passive galaxies
 - Important: deep wide-field imaging and massive spectroscopic surveys
 - ⇒ SuprimeCam (Subaru), VST, VISTA, VIMOS upgrade, FMOS (Subaru)
 - Internal physics and morphologies of galaxies at 1<z<3
 - Important: high spatial resolution and spatially resolved spectroscopy
 - ⇒ HST, NACO, SINFONI, OSIRIS (GTC), MUSE, KMOS, HAWK-I with AO, JWST, E-ELT
- Objects at very high redshifts ('first light')
 - Search for Ly- α emitters, IGM at high z
 - Important: deep surveys, spectroscopic follow-up
 - SuprimeCam (Subaru), X-Shooter, NACO, OSIRIS (GTC), LRIS (Keck), DEIMOS (Keck), HAWK-I with AO, MUSE, KMOS, EMIR (GTC), JWST, E-ELT





Fashions and other transient phenomena ESO top ten cited papers are all supernovae and GRBs - This is more a sign of fashion than sound physics AGNs – topic of the 4m telescopes - Topic for 8m telescopes? Metal-poor stars – originally 8m (e.g. First Stars programme) - And now?





When opportunity knocks

SN 1987A • One in a century object? Comets

Unique objects

Hale-Bopp, Hyakutake, 73P/Schwassmann Wachmann 3, Shoemaker-Levy 9, Halley
Near-Earth objects
Solar system event
Spots on Jupiter
Volcano eruption on Io?
Wesley, 35cm

Role of 2-4m telescopes

- Workhorses of optical/IR astronomy
 - Distributed resource
 - Access at the university level
 - Many national telescopes
 - Access for many astronomers
 - Develop specific strengths
 - E.g. time series, large samples
 - Examples of successful the 4m telescopes over the past decade
 - AAT/2dF, CFHT/Legacy Survey, ESO 3.6m/HARPS, WHT/SAURON and PN.S



Role of 8-10m telescopes

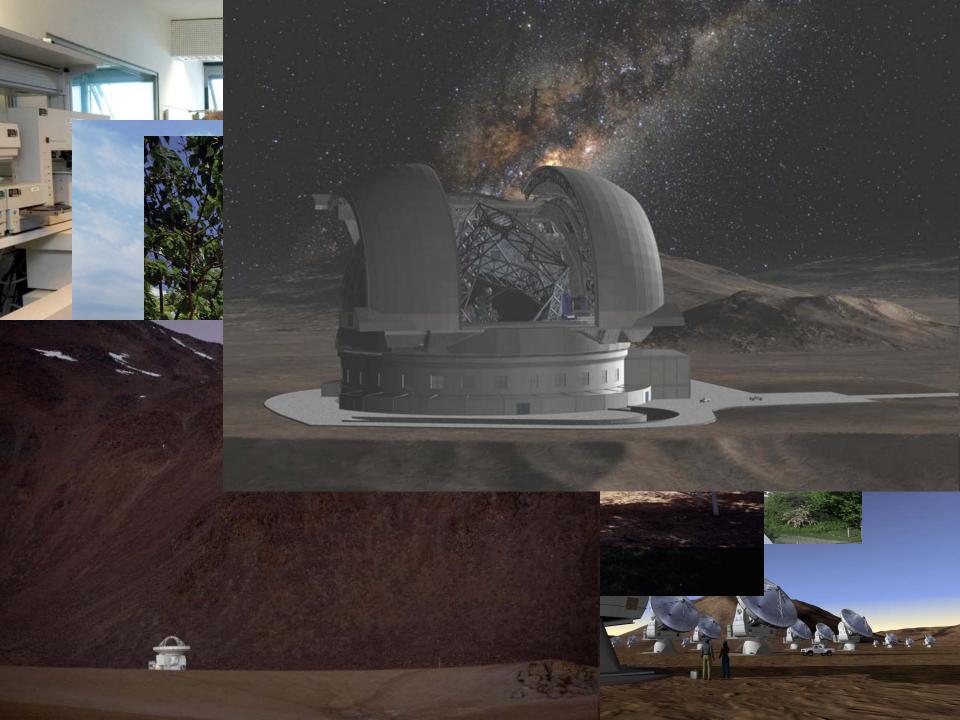
- Workhorses of optical/IR astronomy
 - Distributed resource
 - In general have an open access policy
 - Access for many astronomers across all regions
 - Currently cover all observational parameters
 - In the future develop specific strengths
 - E.g. time series, wide-field spectroscopy



ESO's program

- La Silla Paranal Observatory
 - Continue operations of these southern telescopes
 - Implement second generation instruments (VLT/VLTI)
 - Key surveys with VST and VISTA
 - Long-term programs for unique science on La Silla
 - Prepare for ALMA science with APEX
- Deliver ALMA on time and budget
- Design world-leading E-ELT and secure funding for construction and operations





ESO's world

cching bei

Paranal La Silla Santiago

ESO's sites

Earth at Night More information available at: http://antwrp.gsfc.nasa.gov/apod/ap001127.html Astronomy Picture of the Day 2000 November 27 http://antwrp.gsfc.nasa.gov/apod/astropix.html

La Silla Paranal

Continue operations with new instruments
 FORS2, ISAAC, UVES, FLAMES, NACO, SINFONI, CRIRES, VISIR, HAWK-I, VIMOS, X-Shooter, KMOS, AOF, MUSE, SPHERE
 MIDI, AMBER, PRIMA, GRAVITY, MATISSE

Continue operations with long-term programmes

 Covers sub-mm and mm wavelendths 0.3 to 3 mm
 SHFI (Swedish Heterodyne Facility Instrument), LABOCA, SABOCA, APEX-SZ, CHAMP+, FLASH



La Silla

La Silla

- Medium-size telescopes
 - 3.6m: HARPS for exo-planet searches
 - 3.5m NTT: EFOSC2 & visitor instruments
 - 2.2m in partnership with MPG
- Small telescopes
 - Closed/funded
 externally





La Silla: 5 Operational Instruments



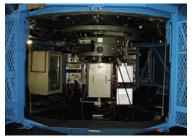




SOFI



EFOSC2

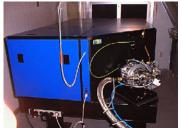


2.2m

WFI



FEROS



GROND





Access to La Silla 3.6m and NTT

- Through the OPTICON Access program non-ESO member state astronomers can get their observing costs reimbursed
- Simply apply to ESO under this program
- So far, this channel has not been used very successfully → very few proposals!



Paranal









VLT Instruments

Instruments Operational on the VLT/I

ANTU ISAAC



KUYEN FLAMES



MELIPAL

VISITOR



VISIR

YEPUN HAWK-I

2xFORS



CRIRES





UVES



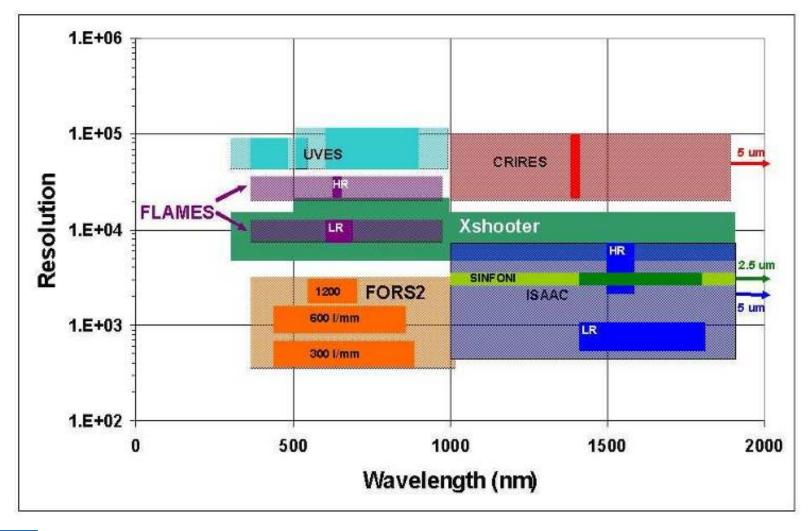
VIMOS



SINFONI



VLT/I Spectral Resolution





Very Large Telescope Interferometer



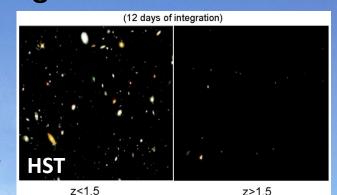


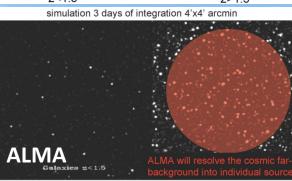


ALMA In Search of our Cosmic Origins

 Science requirements - Detect CO and [CII] in Milky Way galaxy at z=3 in < 24 hr Dust emission, gas kinematics in proto-planetary disks **Resolution to match Hubble**, JWST and 8-10m with AO **Complement to Herschel** pecifications 66 antennas (54x12m, 12x7m) 16 km max baseline (< 10mas) 30-1000 GHz (10-0.3mm),

up to 10 receiver bands





Z>1.5

850 GHz

5AU

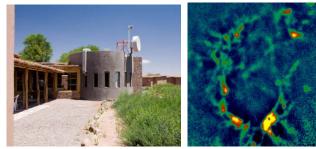


z<1.5

Chajnantor

- APEX
 - 12m sub-millimeter antenna, operated by ESO @ Sequitor
 - MPG (50%), Sweden (23%) and ESO (27%)
- ALMA
 - Transformational science
 - 66 antennas at 5050m
 - Operations support at 2950m
 - Global partnership with North America East Asia & Chile





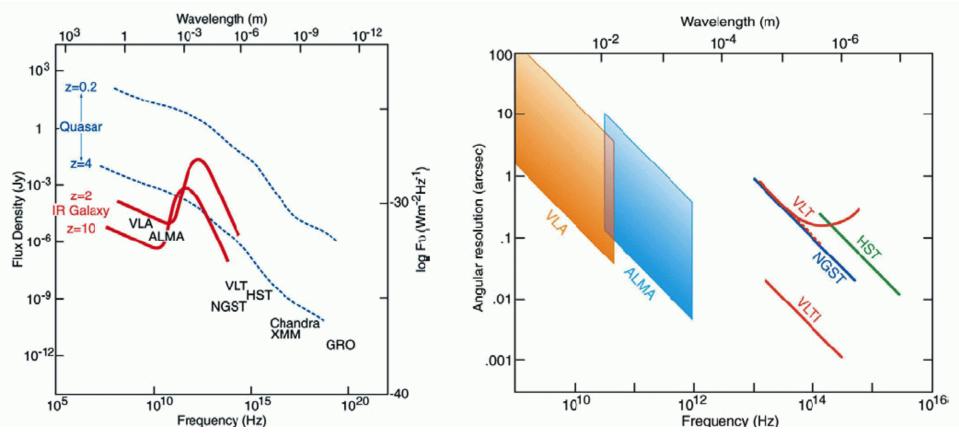




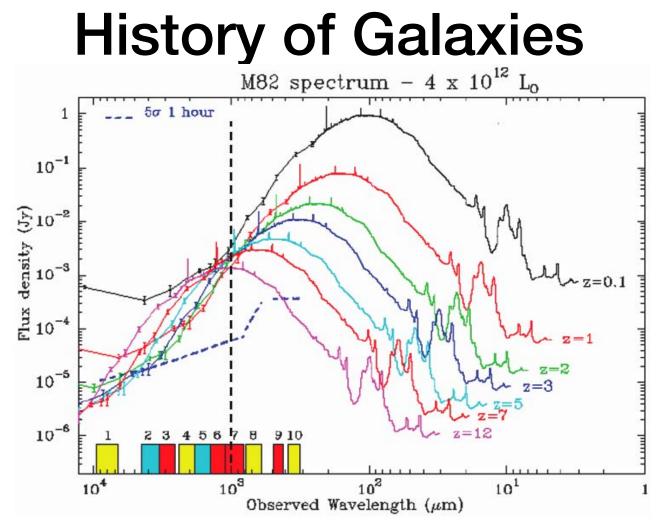
First antenna at 5000m



Sensitivity and Resolution









In the (sub-)millimeter the inverse Kcorrection compensates for the distance as z increases

ALMA Science

- Star Formation, Proto-planets in nearby disks
- Astrochemistry
- Interstellar medium (Galaxy, Local Group)
- High-redshift deep fields
- +130 projects in first 3yrs DRSP 2.1
 - <u>http://www.eso.org</u>/sci/facilities/alma/science/drsp/



E-ELT

Detailed design study

Baseline 42m primary mirror
Adaptive optics built-in
Industry strongly engaged
Study complete in 2010

Project

Builds on *entire* expertise at ESO *and* in the member states

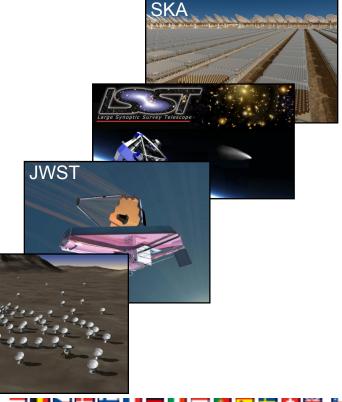
- Construction 2011-2018 - Synergy: JWST/ALMA/SKA



The Science

- Contemporary Science: Exoplanets, planet formation, resolved stellar populations beyond the Local Group, Black Holes, the physics of high redshift galaxies, the expansion of the universe, ...
- Synergies with other top facilities: ALMA, JWST, survey telescopes, SKA, ...
- Discoveries: opening new parameter space in spatial resolution and sensitivity, ...



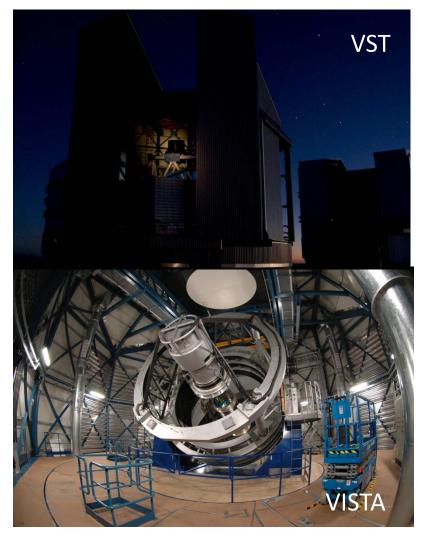




The Survey Telescopes

- Under construction

 VST 2.6m for optical
 Completion in 2010
 - VISTA 4.1m for infrared
 - Currently in science verification
- Science
 - Multi-year program of large public surveys
 - Coordinated by ESO
 - Develops European survey capability





ESO Archive

- The ESO data archive
 - is a rich source of excellent data
 - abstracts of previous proposals available
 - data public one year after they have been delivered to the PI
 - great way to compete with your competitor, if they got observing time
 - easy retrieval and selection of calibration
 data



Get involved

- Participate in ESO activities
 - get to know the organisation better
 - active interactions with ESO people
- Have a lively scientific exchange with the (European) astronomical community
 - conferences, workshops
 - regularly publish your results



European Extremely Large Telescope

